

It's About Time !!!!!

ZITS JERRY SCOTT & JIM BORGMAN



Timing for VLBI

⌘ **Tom Clark**

formerly at NASA Goddard Space Flight Center

⌘ **Rick Hambly**

CNS Systems, Inc.

IVS TOW Meeting

Haystack – May 9-12, 2005

What Timing Performance Does VLBI Need?

- ⌘ The VLBI community (Radio Astronomy and Geodesy) uses Hydrogen Masers at 40-50 remote sites all around the world. To achieve $\sim 10^\circ$ signal coherence for ~ 1000 seconds at 10 GHz we need the two oscillators at the ends of the interferometer to maintain relative stability of $\approx [10^\circ / (360^\circ \cdot 10^{10} \text{Hz} \cdot 10^3 \text{sec})] \approx 2.8 \cdot 10^{-15}$ @ 1000 sec
- ⌘ To correlate data acquired at 16Mb/s, station timing at relative levels ~ 50 nsec or better is needed. After a few days of inactivity, this requires $\approx [50 \cdot 10^{-9} / 10^6 \text{sec}] \approx 5 \cdot 10^{-14}$ @ 10^6 sec
- ⌘ In Geodetic applications, the station clocks are modeled at relative levels ~ 30 psec over a day $\approx [30 \cdot 10^{-12} / 86400 \text{sec}] \approx 3.5 \cdot 10^{-16}$ @ 1 day
- ⌘ Since VLBI defines UT1, we need to control $[\text{UTC}_{(\text{USNO})} - \text{UTC}_{(\text{VLBI})}]$ to an accuracy ~ 100 nsec or better.

The difference between Frequency and Time

Oscillators and Clocks

Oscillator

- Pendulum
- Escapement Wheel
- Crystal Oscillator
- Oscillator Locked to Atomic Transition
 - Rubidium (6.8 GHz)
 - Cesium (9.1 GHz)
 - Hydrogen Maser (1.4 GHz)

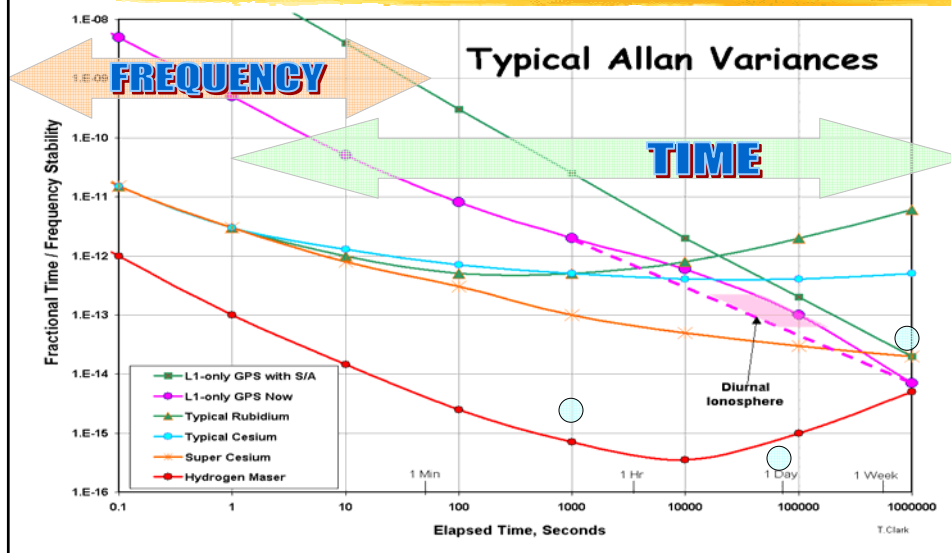
Integrator and Display = Clock

- Gears
- Electronic Counters
- Real Clocks

Events that occur with
a defined
FREQUENCY
nsec -- minutes

Long-Term
TIMING
seconds - years

The Allan Variance – A graphical look at clock performance



Why do we need to worry about “Absolute Time” (i.e. Accuracy) in VLBI?

- To get the correlators to line up for efficient processing, the relative time between stations needs to be known to ~ 100 nsec.
- The correlators maintain their “magic tables” that relates the GPS timing data reported by different stations to each other.
- In the past, geodetic and astronomical VLBI data processing has been done by fitting the data with “station clock polynomials” over a day of observing, and then discarding these results as “nuisance parameters” that are not needed for determining baseline lengths, source structure, etc.
- The uncalibrated and unknown offsets now range from 1-10 usec at many VLBI stations.

Why do we need to worry about “Absolute Time” (i.e. Accuracy) in VLBI?

• The **ONLY** reason for worrying about “absolute time” is to relate the position of the earth to the position of the stars:

- Generating Sidereal Time to point antennas.
- Measuring UT1 (i.e. “Sundial Time”) to see changes due to redistribution of mass in/on the earth over long periods of time.
- Knowing the position of the earth with respect to the moon, planets and even the the GPS satellites.

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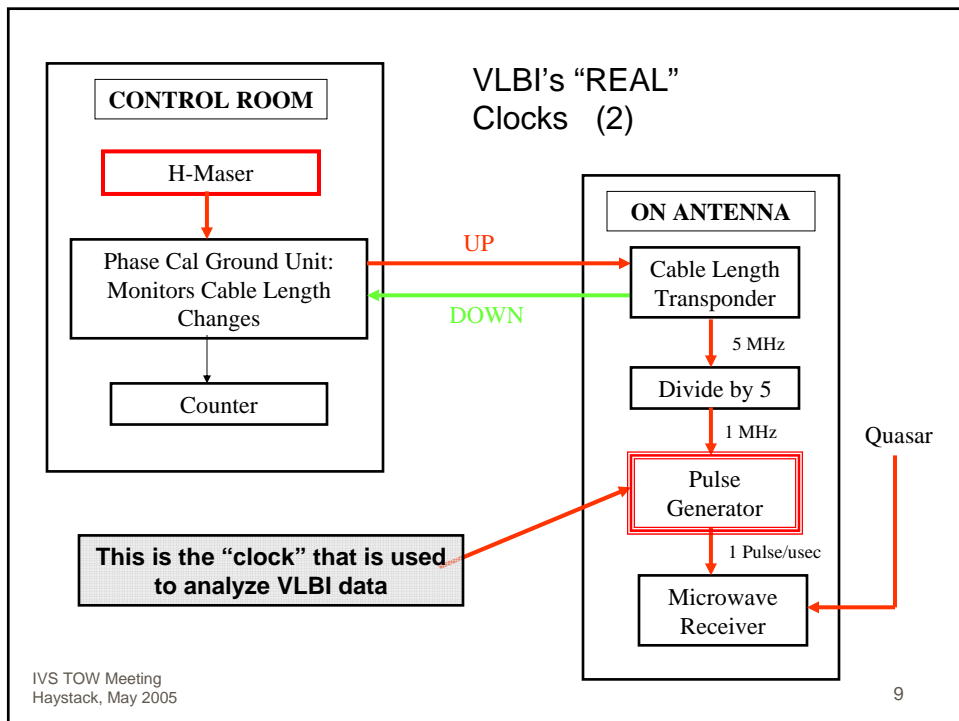
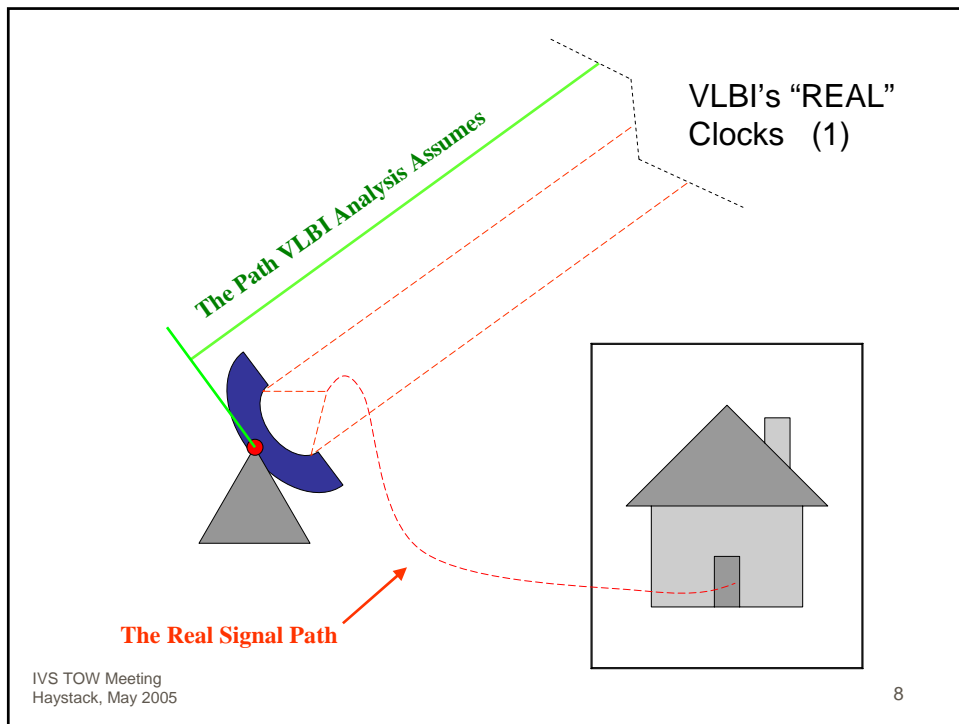
Why do we need to worry about “Absolute Time” (i.e. Accuracy) in VLBI?

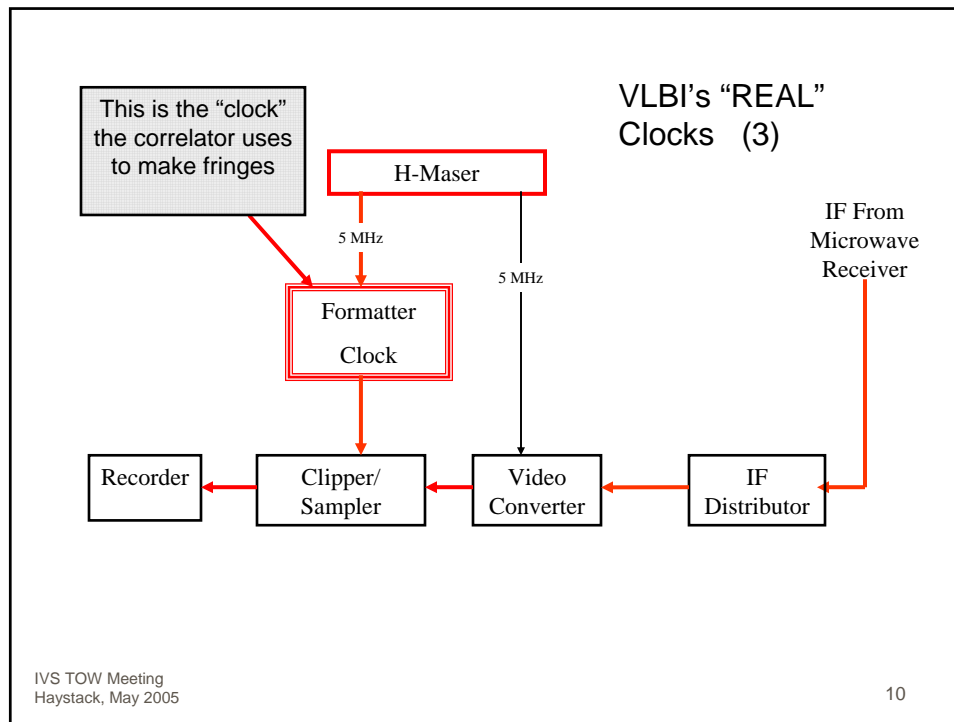
At the stations this means that we will need to pay more attention to timing elements like

- Frequency Standard and Station Timing
- The lengths of cables
- The geometry of the feed/receiver to the antenna.
- Calibration of instrumental delays inside the receiver and backend. The development of new instrumentation is needed.
- The care with which system changes are reported to the correlators and the data analysts.

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Setting VLBI Clocks Time & Rate with GPS -- 3 possible ways--

- ☒ Compare two distant clocks by observing the same GPS satellite(s) at the same time (called Common View)
 - Requires some intervisibility between sites
 - Requires some near-Real-Time communication
 - Links you directly to the "Master Clock" on the other end at ~1 nsec level
- ☒ Use Geodetic GPS receivers (i.e. as an extension of the IGS network)
 - Requires high quality (probably dual frequency) receiver (TurboRogue, Z12, etc), but it's hard to gain access to the internal clock.
 - Requires transferring ~1 Mbyte/day of data from site
 - Requires fairly extensive computations using dual-frequency data to get ~300 psec results with ionosphere corrections
 - Allows Geodetic community to use VLBI Site for geodesy & ionosphere network

👉 **Blindly use the Broadcast GPS Timing Signals as a clock**

- **Single Frequency L1 only (until 2004)**
- **Yields ~10 nsec results with < \$1000 hardware**

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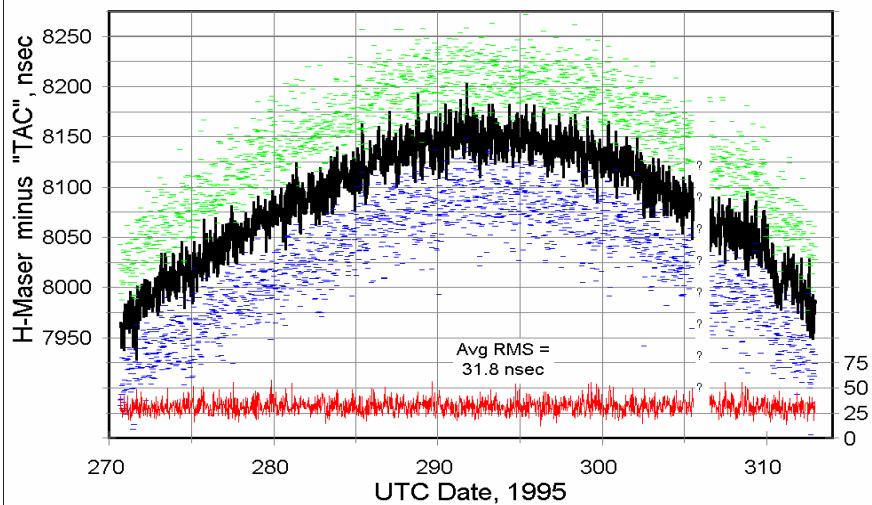
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An Isolated, Remote VLBI Site -- Urumqi in Xinjiang Province, China



An Early Example of "Blind" GPS Timing with a 6 channel receiver

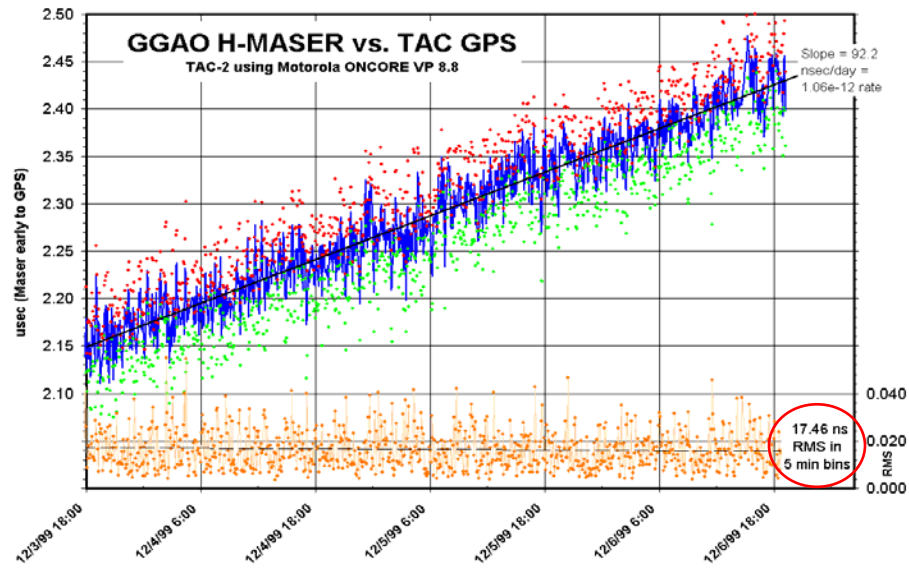
ONSALA H-Maser vs "TAC" GPS



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Before S/A was turned off (8-channel) . . .



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GGAO (Goddard Geophysical & Astronomical Observatory)

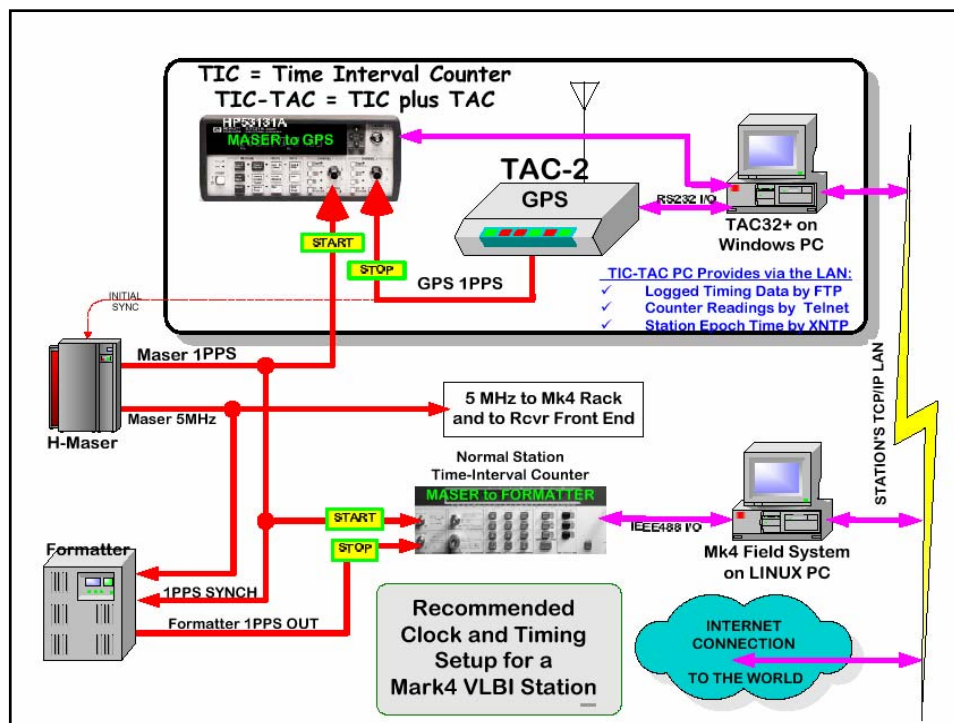


How we got ~30 nsec timing even with S/A

- ⌘ Start with a good timing receiver, like the Motorola ONCORE
- ⌘ Average the positioning data for ~1-2 days to determine the station's coordinates. With S/A on, a 1-2 day average should be good to <5 meters. Or if the site has been accurately surveyed, use the survey values.
- ⌘ Lock the receiver's position in "Zero-D" mode to this average.
- ⌘ Make sure that your Time-Interval Counter (TIC) is triggering cleanly. Start the counter with the 1 PPS signal from the "house" atomic clock and stop with the GPS receiver's 1PPS.
- ⌘ Average the individual one/second TIC reading over ~5 minutes.
- ⌘ **These steps were automated in the SHOWTIME and TAC32Plus Software.**

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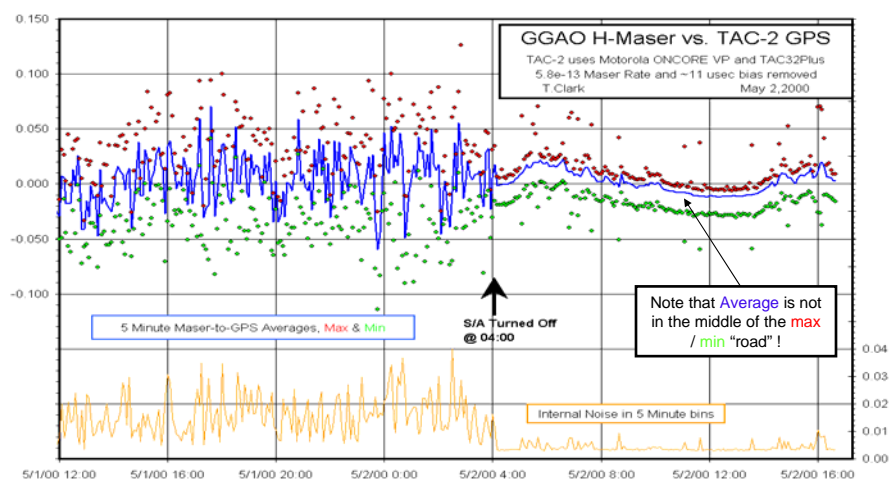


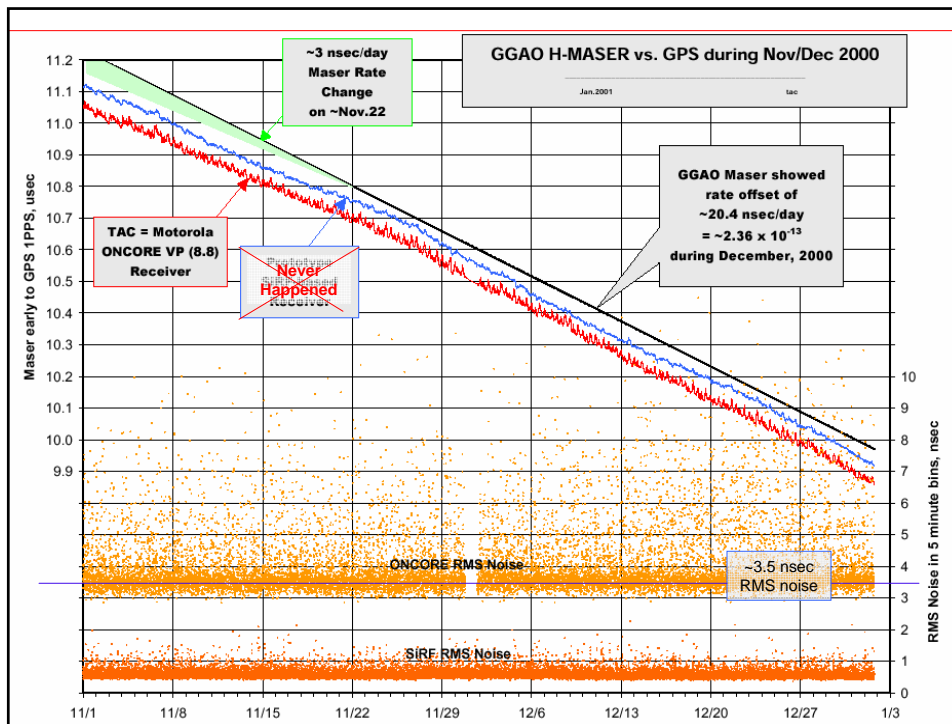
Let Us Now Discuss . . .

- ⌘ What happened when S/A was turned off on May 2, 2000.
- ⌘ Sawtooth and Glitches
- ⌘ Some recent results obtained with Motorola's newest low cost timing receiver (the M12+)

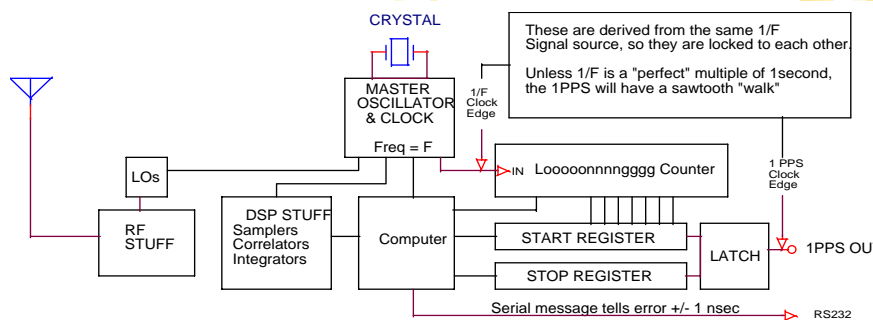
What happened when S/A went away?

Using 8-channel Motorola ONCORE VP Receiver . . .



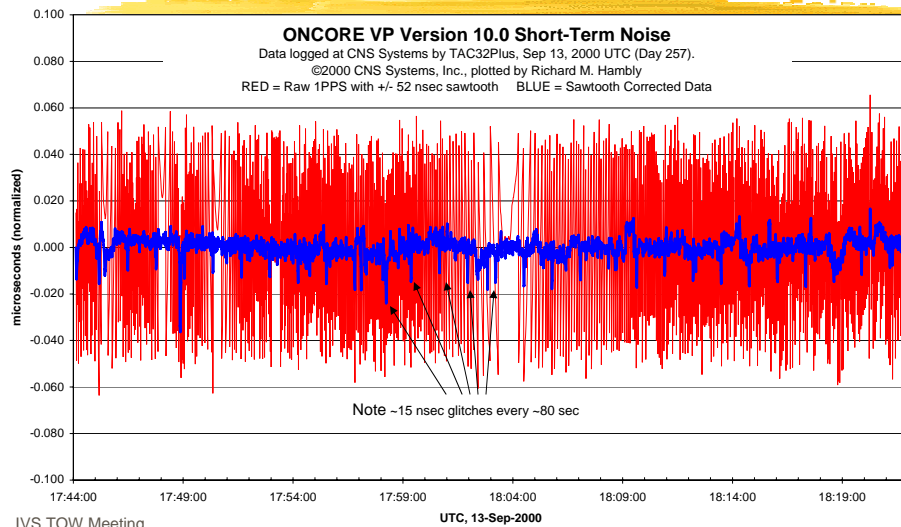


What is the sawtooth effect ????



- For the older Oncore, $F=9.54$ MHz, so the 1/F sawtooth has a range of ± 52 nsec (104 nsec peak-to-peak)
- The new Oncore M12+ has $F \approx 40$ MHz, so the sawtooth has been reduced to ± 13 nsec (26 nsec).

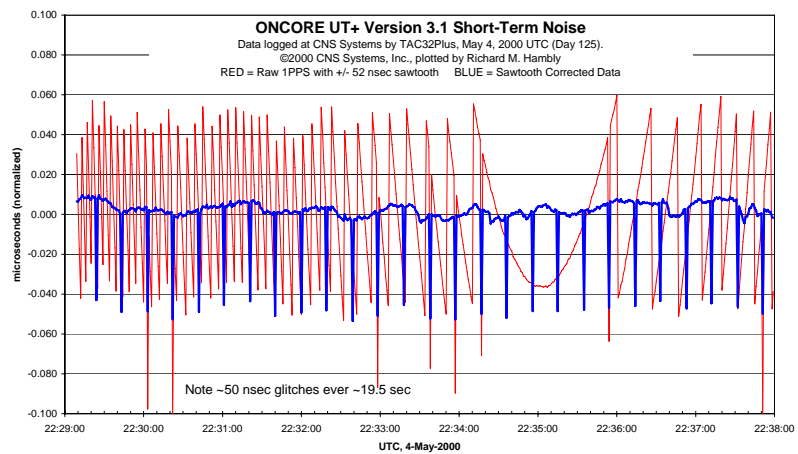
An example of 1PPS sawtooth Motorola VP (10.0)



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An example of 1PPS sawtooth Motorola UT+ (3.1)



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CNS Systems' Test Bed at USNO

Calibrating the "DC" Offset of the new M12+ receiver.

We have observed that the ONCORE firmware evolution from 5.x \Rightarrow 6.x \Rightarrow 8.x \Rightarrow 10.x has been accompanied by about 40 nsec of "DC" timing offsets.

Motorola tasked Rick to make the new M12+ receiver be correct.



Tac32Plus software simultaneously processes data from four Time Interval Counters and four CNS Clocks, writing 12 logs continuously.

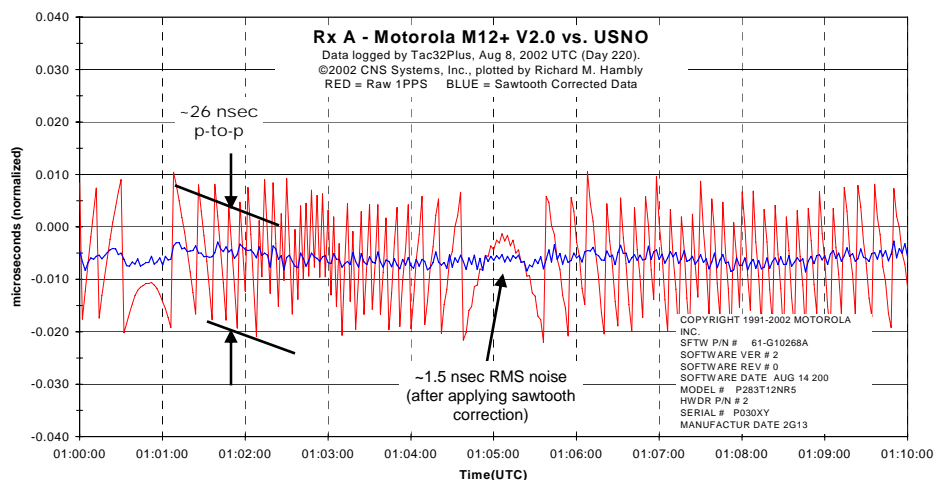


Time Interval Counters compare the 1PPS from each CNS Clock (M12+) against the USNO's UTC time tick.

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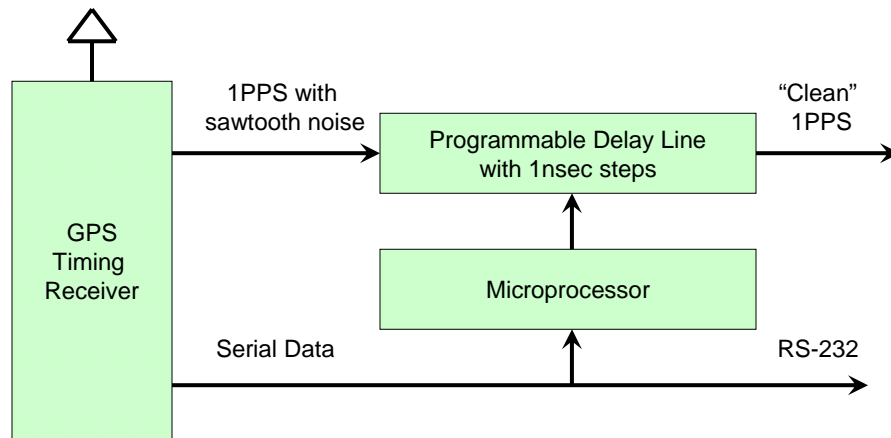
An example of 1PPS sawtooth with the new Motorola M12+ receiver



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How could the sawtooth noise be eliminated ???




In 2003 we showed this potential solution ...

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The Future is here now! The CNS Clock II

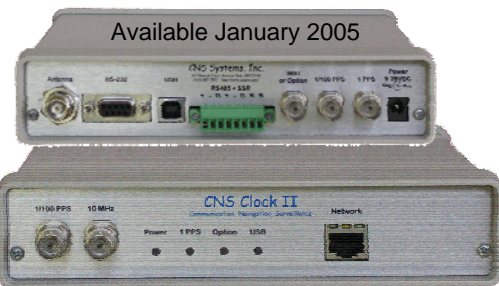
1994 - 2004



1PPS Sawtooth
Correction
Option →



Available January 2005

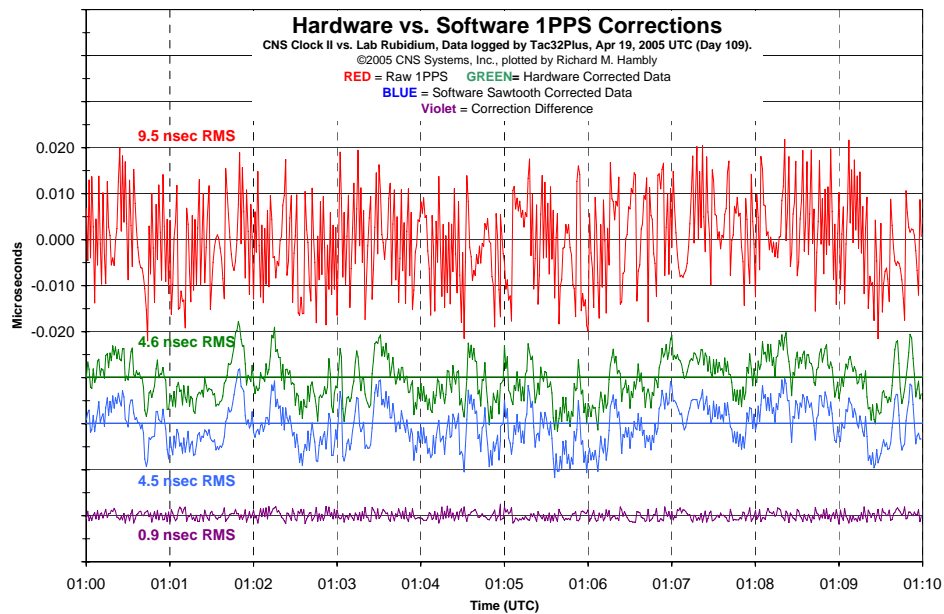
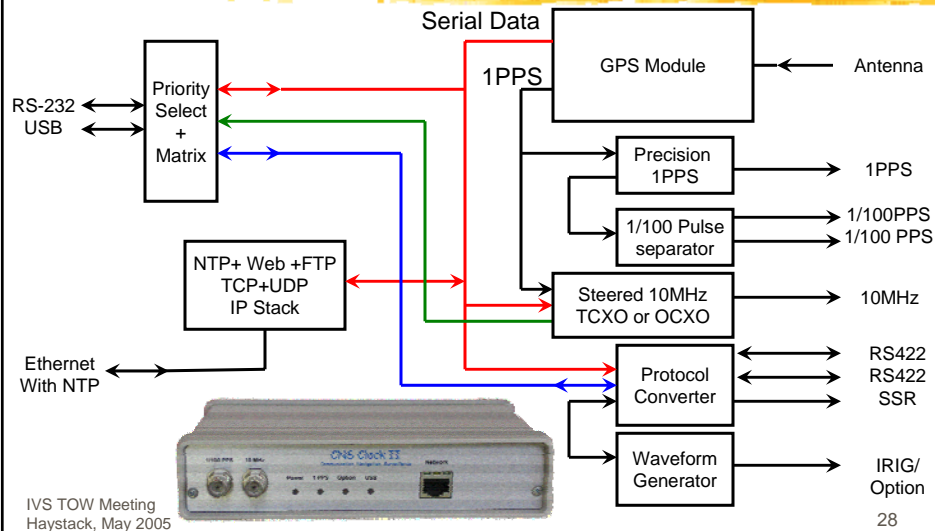


RS-232 Serial Port
 USB 2.0 Port
 TNC(F) Antenna Connector
 Buffered 1 PPS output
 Two buffered 1/100 PPS outputs
 10 MHz output
 2 bidirectional RS-485 ports
 Bipolar (AC/DC) solid state relay out
 Power 9-30 volts @ 500ma
 Options:
 Tx Sequencer output.
 IRIG-B output (modulated, PWM or Manchester).

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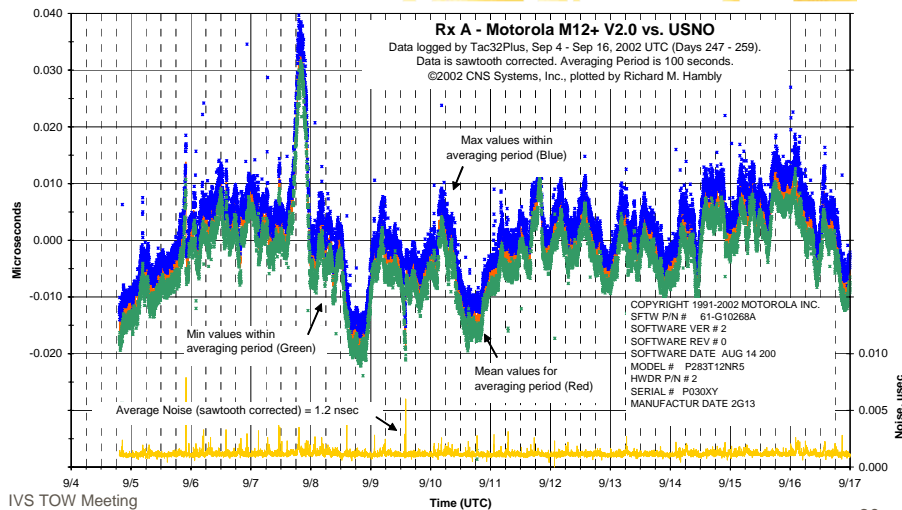
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CNS Clock II Block Diagram

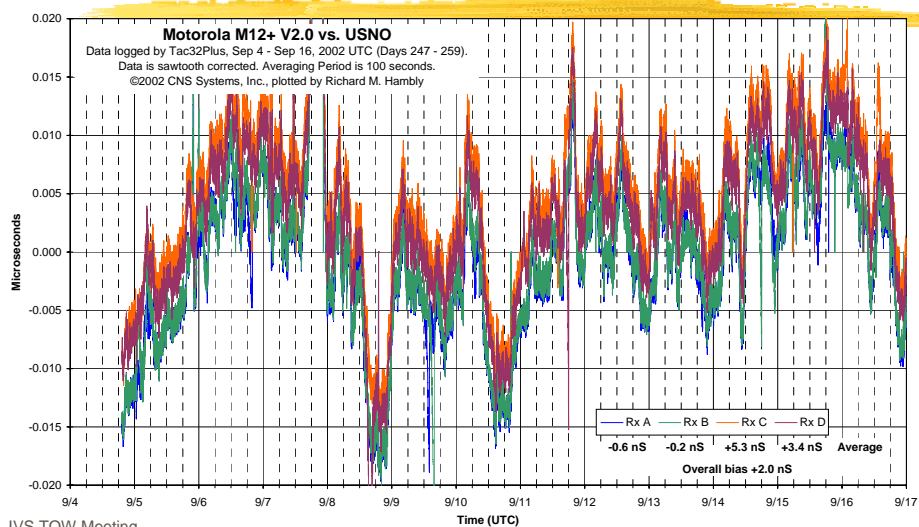


Individual M12 Clock Performance

Receiver (A) average "DC" offset = -0.6 ns



Comparing four M12+ Timing Receivers



What Happened on 9/7/02 ?



September 7, 2002.

This picture is a two hour composite of 85 different photos spanning 21:07 thru 23:10 EDT on Sept. 7th (01:07 thru 03:10 UTC Sep. 8).



September 8, 2002.

This picture is a four hour composite of 140 different photos spanning 20:00 thru 24:00 EDT on Sept. 8th (00:00 thru 04:00 UTC Sep. 9).

Each picture was an 87 second exposure with 3 seconds between frames. The trails on the picture are all due to airplanes. The bright loop is from a plane on final approach into BWI airport. Camera = Canon D60 shooting Hi Resolution JPEG at ISO 100 with TC-80 timer. Lens = Sigma f/2.8 20-40 mm set to 20 mm @ f/4.5

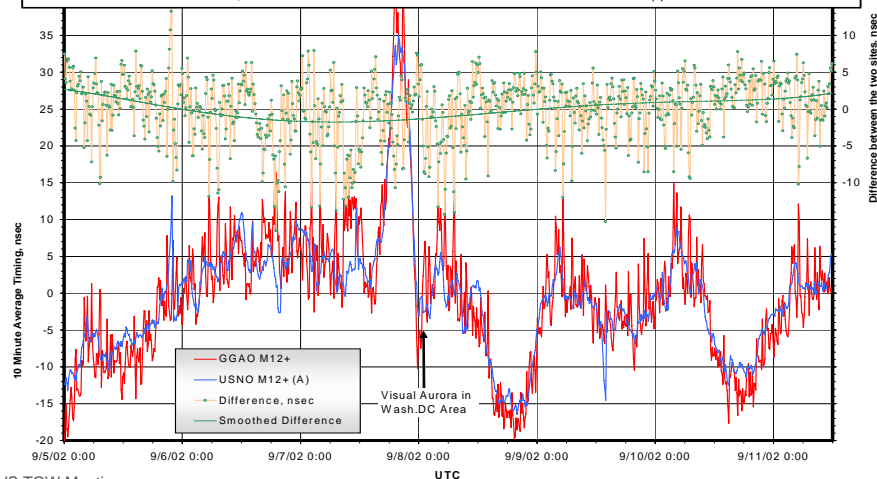
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Short Baseline Test (USNO to NASA GGAO)

Comparing two new Motorola M12+ GPS Timing Receivers over the 21.5 km baseline between the US Naval Observatory (USNO) and the NASA Goddard Geophysical & Astronomical Observatory (GGAO).

Both data sets compare the GPS timing receiver to a local Hydrogen Maser clock.
On both, a linear fit to remove constant clock offset and drift has been applied.



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Where to get information?

These Slides and related material:

<http://gpstime.com>

Information on the CNS Clock and the CNS Clock II:

<http://www.cnssys.com>

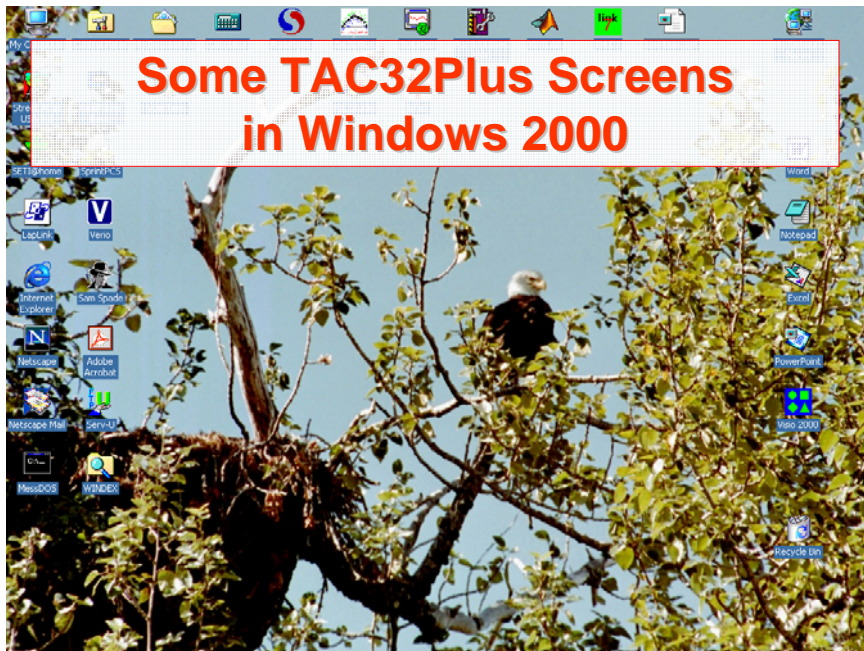
For ONCORE/TAC-2 receiver used as a LINUX xntp server:

<http://gpstime.com>

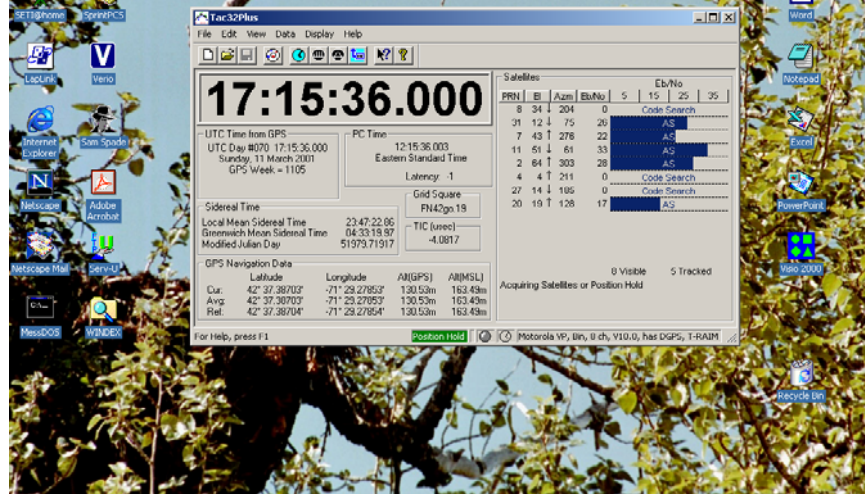
To contact Tom: <mailto:w3iwi@toad.net>

To contact Rick: <mailto:rick@cnssys.com>

Some TAC32Plus Screens in Windows 2000



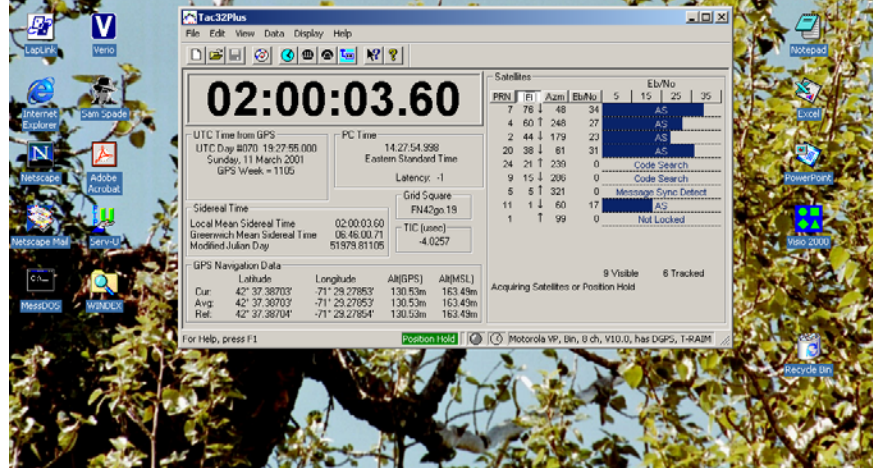
TAC32Plus: DISPLAYS UTC TIME



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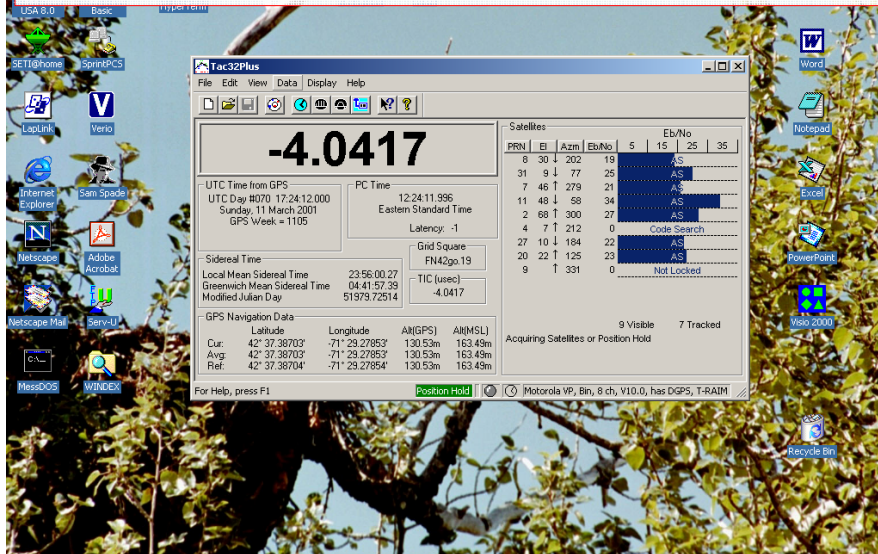
TAC32Plus: DISPLAYS Local Station Sidereal Time (LMST)



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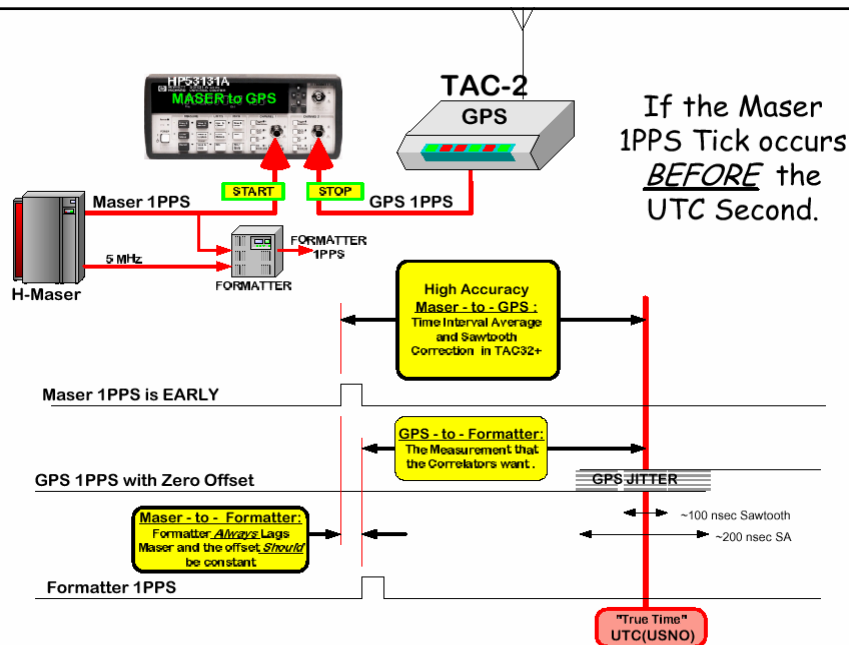
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TAC32Plus: DISPLAYING TIME-INTERVAL COUNTER READINGS WITH CORRECTIONS



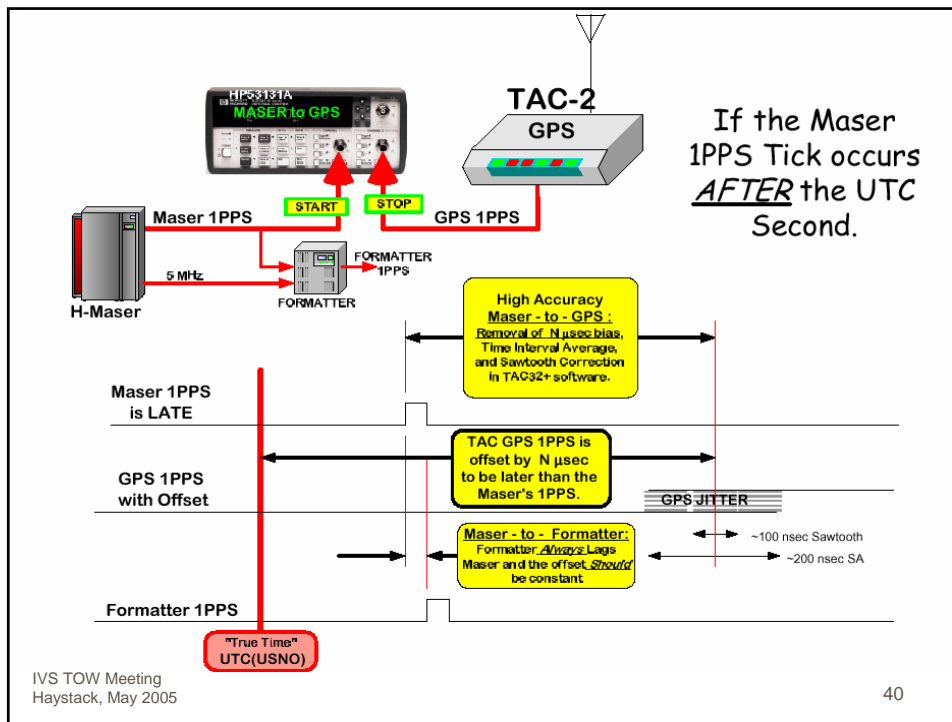
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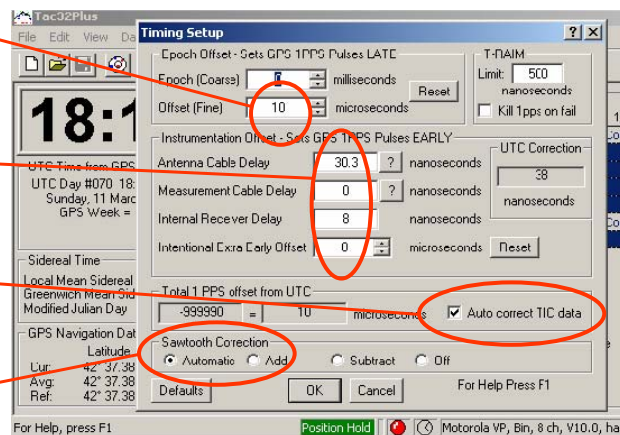
To Make Sure TAC32 is Logging the "true" Maser-to-GPS Time Interval:

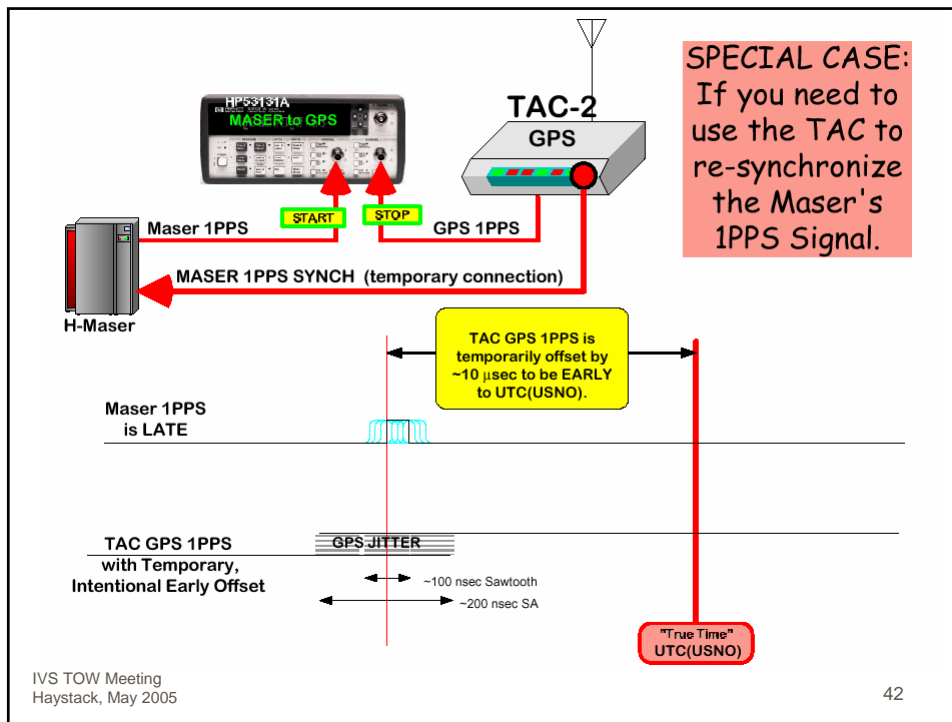
Offset GPS LATE if needed to be certain that GPS 1PPS is later than Maser 1PPS.

Be certain to account for the lengths of all coax cables.

Allow the software to correct for all timing offsets.

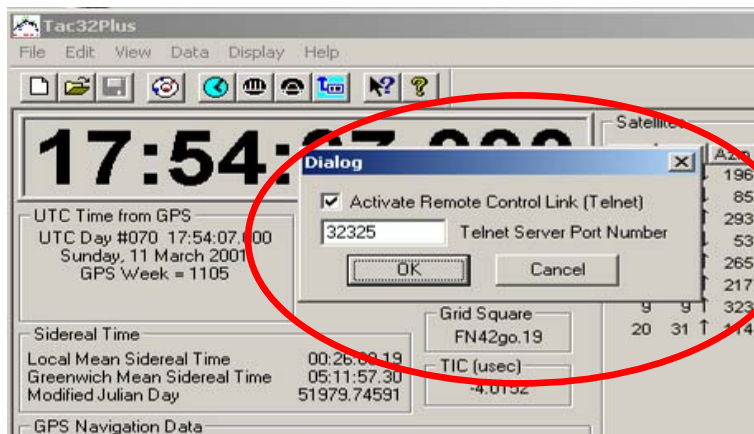
Allow software to correct the 1PPS pulse-to-pulse jitter



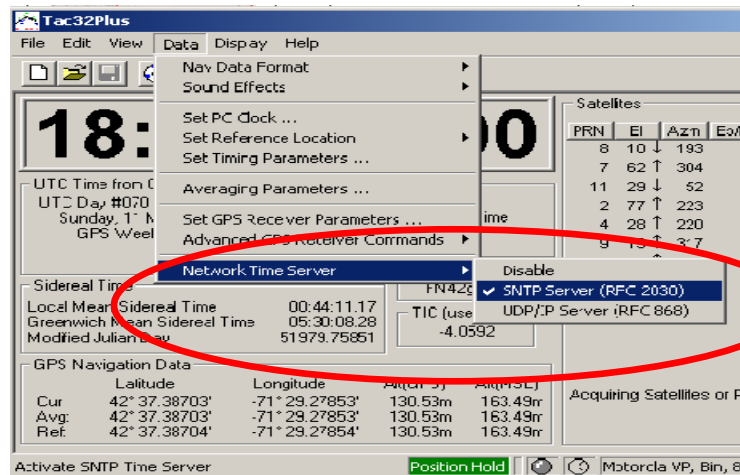


To Activate the LAN Telnet Link between TAC32Plus and the LINUX PC Field System, Hit Control-T:

Then Click on the check-box and the OK button



To Use TAC32Plus as your Station's SNTP Network Timer Server:



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APPENDIX B TEXT MATERIAL

- Field System Documentation
for tacd
- Ed Himwich's documentation
for gpsoff
- David Holland's HP53131
setup notes

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