

# Seshan VLBI Station Report for 2002

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## Abstract

The Sheshan (also called Seshan) 25-meter radio telescope is an alt-az antenna run by Shanghai Astronomical Observatory (SHAO), Chinese Academy of Sciences (CAS). It is one of the five main astronomical facilities of Chinese National Astronomical Observatories. The VLBI station is a member of the EVN, IVS, and APT. We give a short report of the current status and future plans of Seshan VLBI station of Shanghai Astronomical Observatory as an IVS Network station.

## 1. Introduction

The telescope is located about 30 km west of Shanghai. The radio telescope started its operation in 1987.

Station Location: Longitude:  $121^{\circ} 11' 59''$  E; Latitude:  $31^{\circ} 05' 57''$  N.  
Height above sea level: 5 meters (ground).

It is one of the five main astronomical facilities of Chinese National Astronomical Observations. The VLBI station is a member of the EVN, APT and IVS. There is a two-station MKIV data processor and an analysis center of IERS of various space geodetic observations in Shanghai observatory.

## 2. Facilities

### 2.1. Antenna

Diameter : 25 meters

Antenna type: Kashegelun Beam wave-guide

Seat-rack type: Azimuth-pitching ring

Main surface precision: 0.65 mm (rms)

Point precision:  $20''$  (rms)

Rolling range: Azimuth :  $-86^{\circ} - 425^{\circ}$ ; Elevation:  $5^{\circ} - 88^{\circ}$

Maximum rolling speed: Azimuth :  $0.55^{\circ}/\text{sec}$

Elevation:  $0.28^{\circ}/\text{sec}$

### 2.2. Receiver

Five bands for VLBI observations are available at Sheshan VLBI station: L band (18 cm), C band (6 cm), K band (1.3 cm), and S/X band (13/3.6 cm). The parameters of the receivers are listed in Table 1. Column 1 gives the observation band. The frequency range is listed in column 2, followed by the efficiency of each band in column 3. The receiver type, system temperature, and polarization model are listed in columns 4, 5 and 6, respectively.

The L, C, and K bands are used for astrophysics and S/X double frequencies are used for geodesy. X band is also used for astrophysical observations sometimes.

Table 1. VLBI Receivers of Seshan Satation

| <b>Band</b><br>(cm)<br>(1) | <b>Bandwidth</b><br>(MHz)<br>(2) | <b>Efficiency</b><br>(%)<br>(3) | <b>Type</b><br>(4) | $T_{system}$<br>(K)<br>(5) | Polarization<br>(6) |
|----------------------------|----------------------------------|---------------------------------|--------------------|----------------------------|---------------------|
| 18                         | 1620-1680                        | 40                              | Room Temperature   | ~ 100                      | LCP & RCP           |
| 13                         | 2150-2350                        | 45                              | Room Temperature   | ~ 100                      | RCP                 |
| 6                          | 4700-5100                        | 58                              | Cryogenic          | 45-50                      | LCP                 |
| 3.6                        | 8200-9000                        | 48                              | Cryogenic          | ~ 50                       | RCP                 |
| 1.3                        | 22100-22600                      | ~20                             | Cryogenic          | ~110                       | RCP & LCP           |

### 2.3. Recording System

VLBA, MKIV and S2 recording systems are available now at Seshan VLBI station. MKIV upgrade of Seshan station has completed in 2000. Seshan station has participated in the two-head recording test with 512 Mbit/s organized by EVN in Oct. 2001. Fringes were found to Seshan, both head-stacks successful. The performance of the observing system of Shanghai station has been improved over the last few years.

The Field System has been upgraded to 9.5.17 version and it works well for Seshan station in the second half of 2002. The MKIV recording system works well for EVN and IVS observations, and S2 recording system work well for VSOP observation. Two head stacks recording system has been tested successfully and good fringes have been found to Seshan station.

### 3. Personnel

There are some changes of the staff in Seshan station. The main staff members at Seshan VLBI Station are listed in following Table 2.

Table 2 - The main staff in Seshan VLBI Station

| <b>Name</b>     | <b>Position</b> | <b>Working area</b> | <b>email address</b>      |
|-----------------|-----------------|---------------------|---------------------------|
| Xiaoyu Hong     | Professor       | Head of station     | xhong@center.shao.ac.cn   |
| Wenren Wei      | Professor       | Chief Engineer      | wwr@center.shao.ac.cn     |
| Shi-guang Liang | Professor       | Microwave           | sgliang@center.shao.ac.cn |
| Zhihan Qian     | Professor       | Advisor             | qzh@center.shao.ac.cn     |
| Xinyong Huang   | Senior Engineer | VLBI friend         | xhuang@center.shao.ac.cn  |
| Zhuhe Xue       | Senior Engineer | Terminal software   | zhxue@center.shao.ac.cn   |
| Jiazheng He     | Senior Engineer | Antenna control     | jzhe@center.shao.ac.cn    |
| Qing-yuan Fan   | Senior Engineer | Antenna control     | qyfan@center.shao.ac.cn   |
| Song-lin Chen   | Engineer        | Microware           | slchen@center.shao.ac.cn  |
| Bin Li          | Engineer        | Microware           | bing@center.shao.ac.cn    |
| Jinqing Wang    | Engineer        | operator            | jqwang@center.shao.ac.cn  |
| Huihua Li       | Engineer        | operator            | hhlee@center.shao.ac.cn   |
| Lingling Wang   | Engineer        | operator            | llwang@center.shao.ac.cn  |

#### **4. Current Status and Activities**

A rapid frequency switching system is being restructured in June 2002. The mechanical parts have been completed. The re-equipment of the last mirror has worked well for the S/X and C bands observations at present.

Sixteen geodetic experiments have been run by Seshan Station in 2002. And 24 geodetic experiments will be operated during 2003.

#### **5. Future Plans**

A new Hydrogen Maser Clock has been ordered from Datum for Seshan VLBI station. We expect to receive it by the end of 2003.